

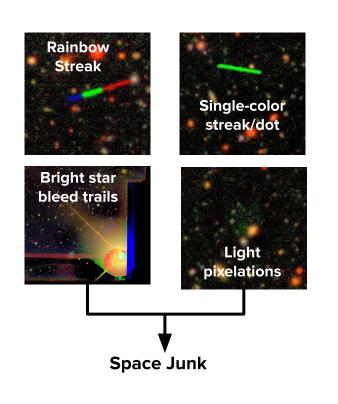
# Finding Space Junk with the World's Biggest Telescopes

Rebecca Bell, Pavani Jairam, Jiayue Xu



### What is Space Junk?

Astronomers use images of deep space to study the universe. These images are polluted with "space junk" that obstructs their view.

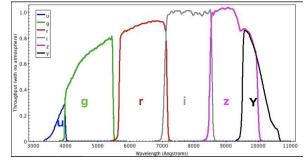


### **Data: The Dark Energy Survey**

### **DES Data**

- International effort to map galaxies, detect supernovae, and reveal patterns of dark energy
- 6 deep fields, 363 4200x2200 images

### Each image consists of 6 bands

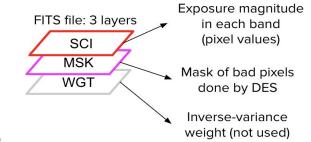


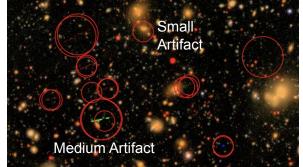
### **Manual Mask Data**

- Artifact location identified by volunteers
- 231 images contain small and medium masks

Approach 1:

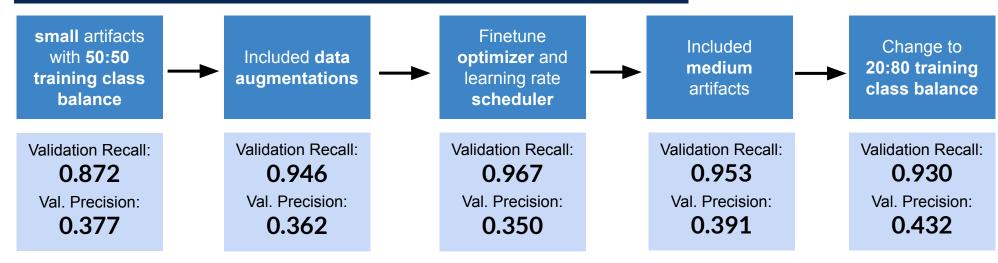
### Each band stored in FITS file

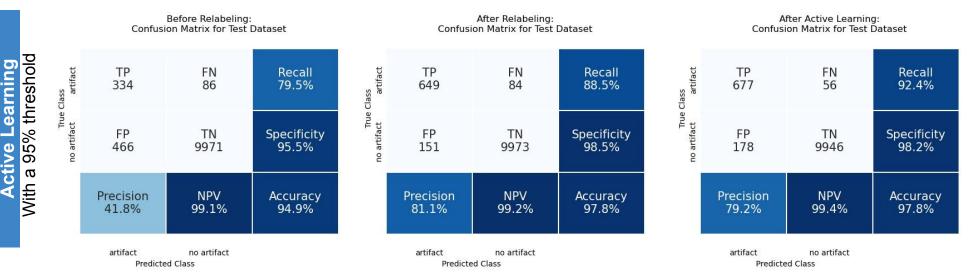




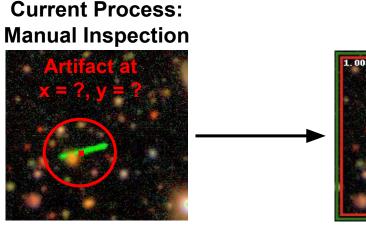
Each circle mask centered at x,y coordinates of artifacts

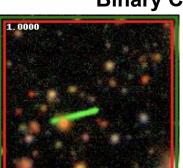
# **Binary Classification: Methods & Results**



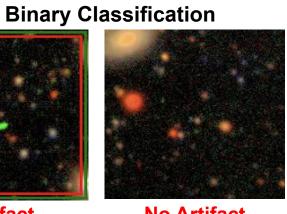


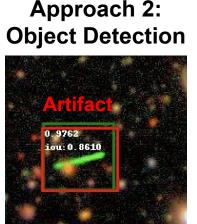
# Deep Convolutional Models to Find Space Junk





**Artifact** 



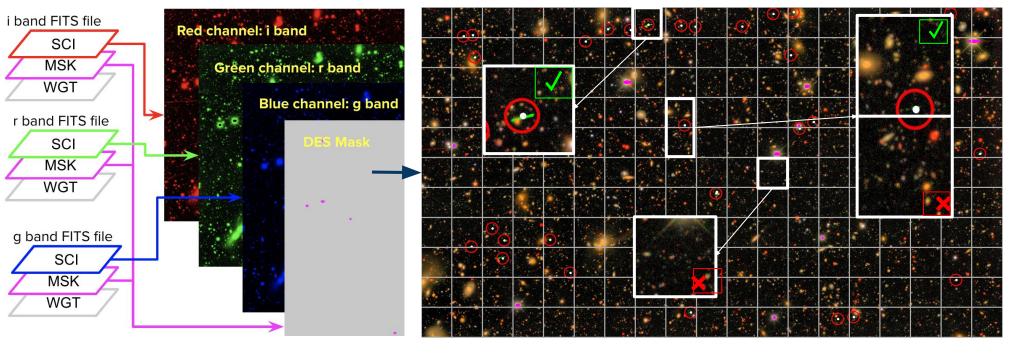


**No Artifact** 

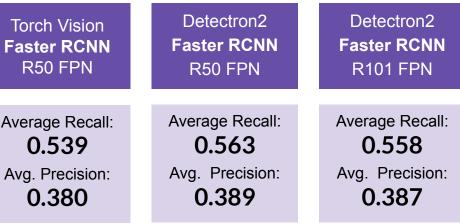
Manually detecting space junk is time consuming, inefficient, and leaves room for human error. Our mission was to build an algorithm that would automate this detection process.

### **Data Pre-Processing**

- To transform FITS files into JPEGs, we extracted the SCI layer of the i, r, g-bands and mapped them to red, green, and blue channels. Then, each of the MSK layers are combined into a single array, which, when applied to the RGB array, shows magenta colored masks.
- We located the artifacts using x,y coordinates. After dividing the image into tiles (200x200 for classification, 600x600 for detection), we classified each by the percentage of tile that is masked.



### **Object Detection: Methods & Results**



**Binary Classification** 

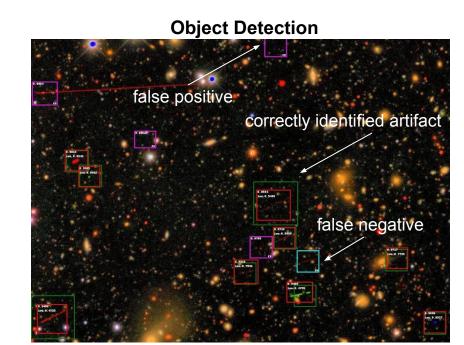
Detectron2 **Faster RCNN** X101 FPN Average Recall: 0.575 Avg. Precision: 0.407

Detectron2 Retinanet R50 FPN Average Recall: 0.582 Avg. Precision: 0.391

Detectron2 Retinanet R101 FPN Average Recall: 0.602 Avg. Precision: 0.399

## **Comparison of Results**

# correctly identified



When the results of our two models are stitched together, they generate the full-sized images above.

# Next Steps

false negative

We plan on continuing our work into the semester. We hope to extend our models into all 6 channels, apply active learning to object detection, fine tune, and package the software for DES researchers.

### **Acknowledgements**

We would like to thank our project manager Kevin Liang, and project leads Prof. Michael Troxel, Prof. Daniel Scolnic, Prof. Christopher Walter, and Bruno Sanchez for their support.